

Ground Based Low Redshift Supernova Sample

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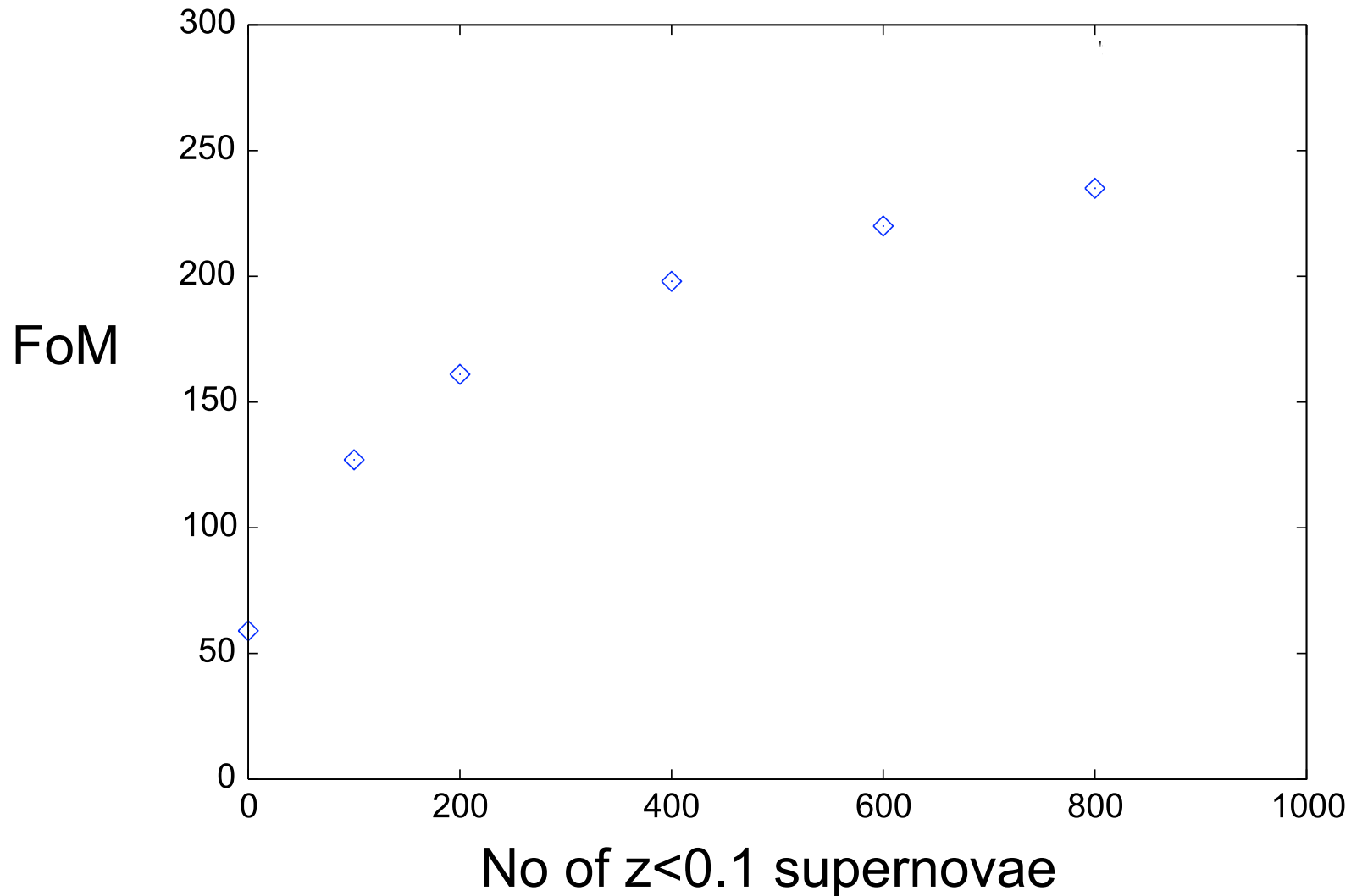
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Need Ground based Low Redshift Sample

- High redshift Supernovs very faint, light redshifted into the infrared
 - Need Space, survey 10's of sq degrees for 6 months
- Low redshift ($z < 0.1$) supernova very rare (volume effect) but bright, light is in the optical
 - Better done from the ground with 1 to 2 m telescopes, $\sim 10,000$ sq degrees over many years

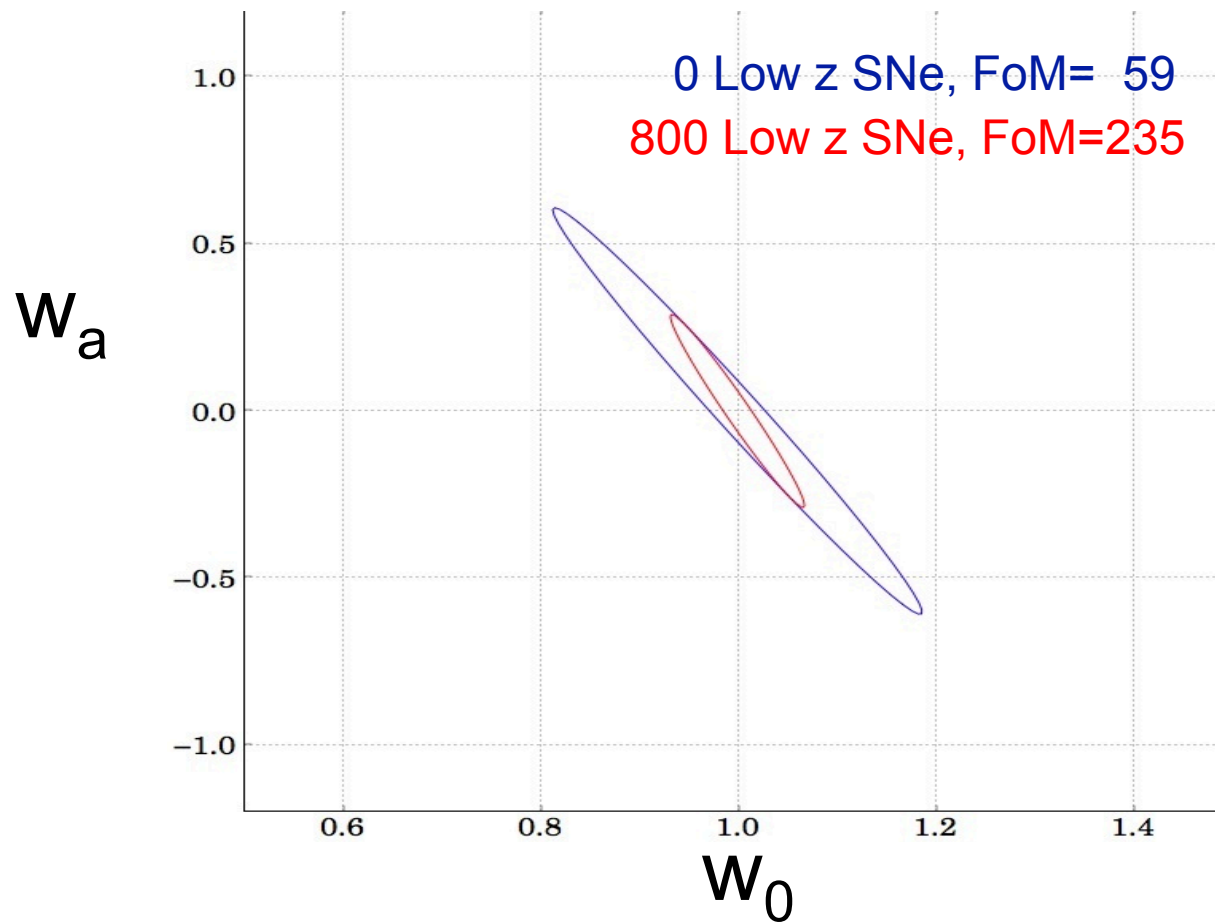
WFIRST Figure of Merit vs no. of low z SNe

800 to 1000 supernovae are important



WFIRST Error ellipses in w_a - w_0

Supernova only, without Stage III Priors



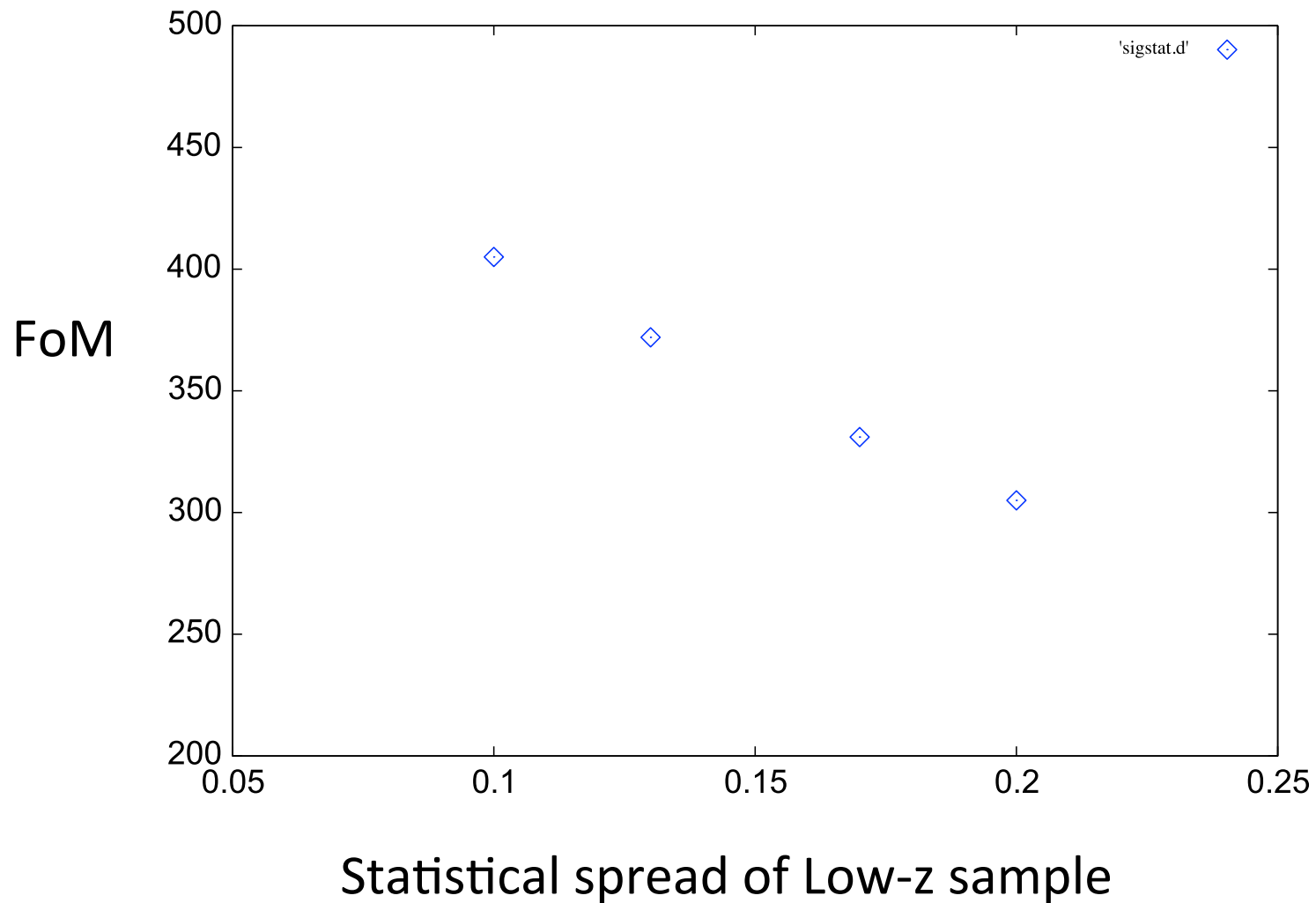
Low Redshift Supernova Surveys

| Survey | Feature | Published | Expected |
|-----------------------------------|--------------------|-----------|----------|
| Collected Present Total Sample | photometric | 128 | |
| SNfactory | spectrophotometric | | 300 |
| CSP-II | near infrared | | 100 |
| PTF | photometric | | 300 |
| LaSilla/QUEST | photometric | | 300 |
| SkyMapper | photometric | | 200? |

Size of Ground Based Low-z sample Probably OK

- WFIRST benefits greatly by 800 or more ground-based low-z supernovae
- Surveys in progress promise ~1300 low-z SNe in the next two or three years
- Promises don't always show up on Hubble diagrams, but can continue surveys longer if necessary
- **BUT *quality of sample is important!!***
And even more so for the control of systematics such as SN population drift.

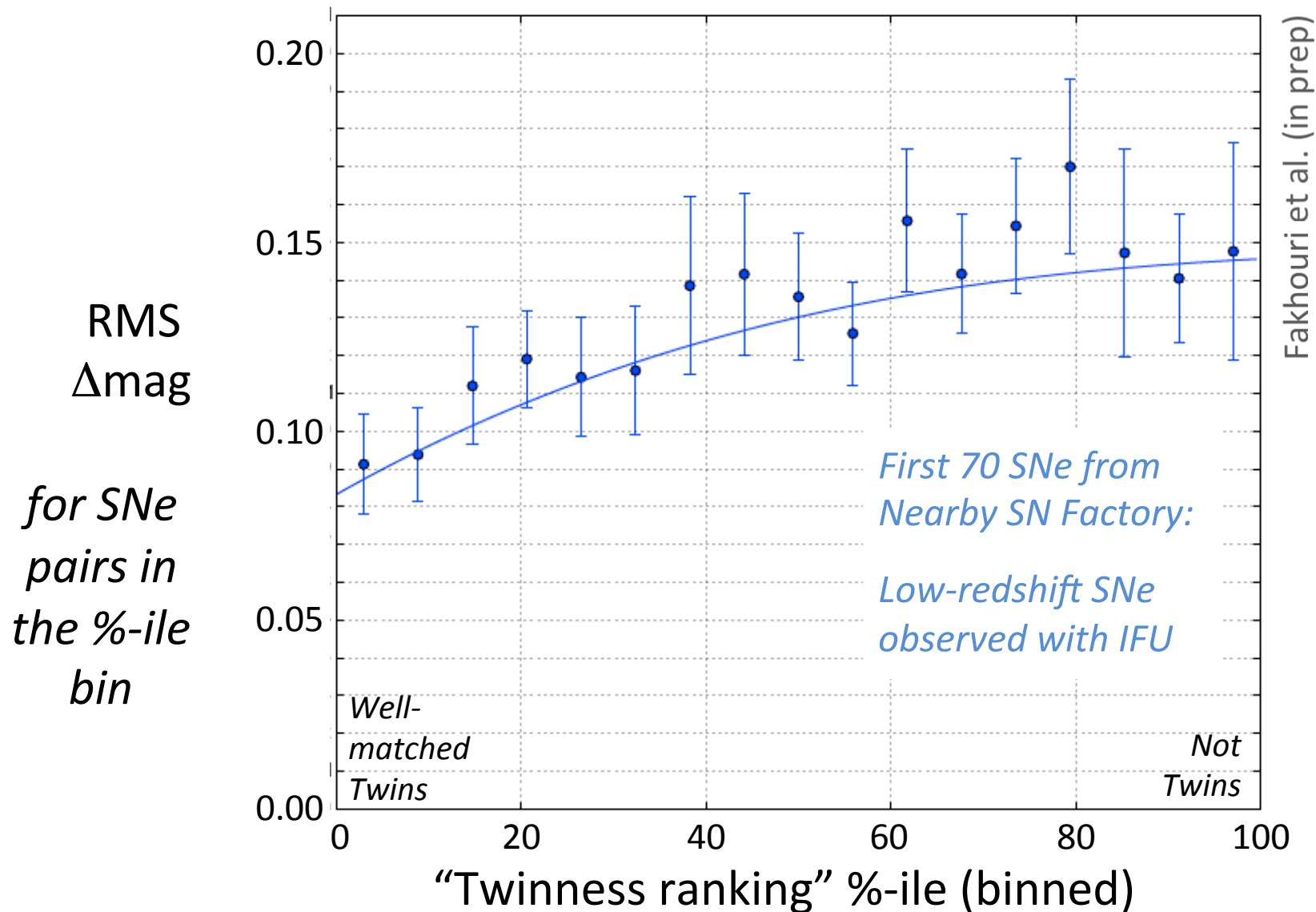
FoM Dependence on Statistical Spread of the Low-z Sample



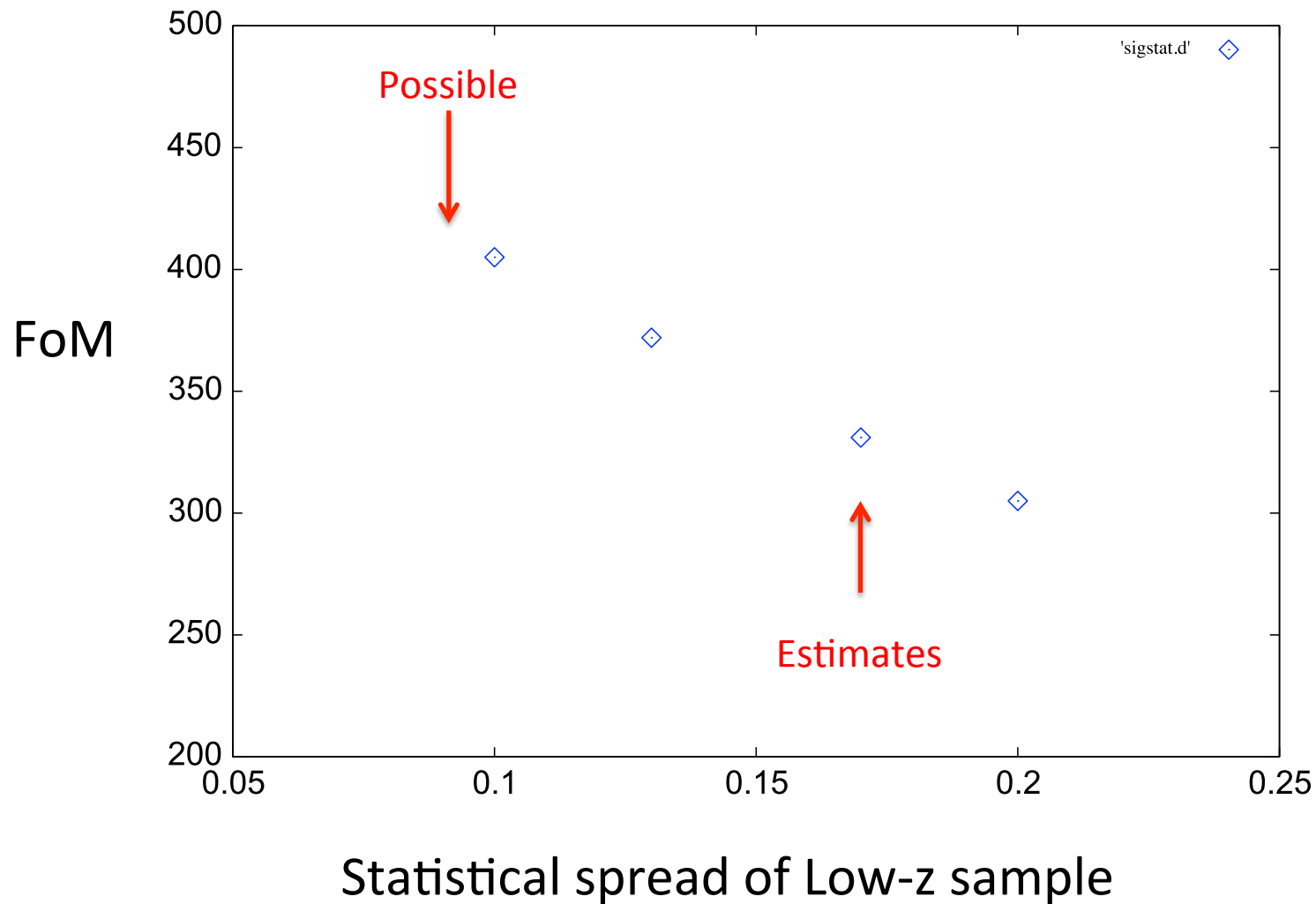
Quality of Samples

- In our previous generations of WFIRST estimates we assumed 800 low z ($z < 0.1$) supernova with a 17% spread in magnitudes and a 1% relative calibration to the WFIRST high z sample
- 17% spread is usual for samples with photometric light curves.
- Matching SNe using spectrophotometric light curves, as from the SNfactory, now achieving a 9% spread (see next slide), better than we assumed in our previous calculations.
- These matching techniques, applied across redshifts (and between ground- and space-based samples), using spectrophotometry with IFU's, makes it possible to constrain the SN population-drift systematic uncertainties.
- Should continue increasing the spectrophotometric ground based sample (a 20% increase in the FoM possible)

Much lower dispersion for spectrophotometrically matched SNe



FoM Dependence on Statistical Spread of the Low-z Sample



Conclusion

- We are well on our way to a sufficiently large sample of ground-based low-redshift supernovae with the quality we assumed in our WFIRST calculations, however...
- ...WFIRST would further benefit significantly from a larger spectrophotometric low-redshift sample, to allow matching the low and high-redshift samples and use the much smaller dispersion of matched samples.

Projects to continue to collect larger spectrophotometric low-z samples are now being developed

- Matching techniques between ground and space based samples, using spectrophotometry with IFU's, also makes it possible to constrain systematic uncertainties, SN population drifts in particular.
- Good calibration is also crucial !!